

Generalized dyons and magnetic dipoles: The issue of angular momentumFrancisco Navarro-Lérida,¹ Eugen Radu,² and D. H. Tchrakian^{3,4}¹*Departamento de Física Atómica, Molecular y Nuclear, Ciencias Físicas, Universidad Complutense de Madrid, E-28040 Madrid, Spain*²*Departamento de Física da Universidade de Aveiro and I3N, Campus de Santiago, 3810-183 Aveiro, Portugal*³*School of Theoretical Physics, Dublin Institute for Advanced Studies, 10 Burlington Road, Dublin 4, Ireland*⁴*Department of Computer Science, Maynooth University, Maynooth, Ireland*
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It is known that a non-Abelian magnetic monopole cannot rotate globally (although it may possess a nonzero angular momentum density). At the same time, the total angular momentum of a magnetic dipole equals the electric charge. In this work we question the generality of these results by considering a number of generalizations of the Georgi-Glashow model. We study two different types of finite energy, regular configurations: solutions with net magnetic charge and monopole-antimonopole pairs with zero net magnetic charge. These configurations are endowed with an electric charge and carry also a nonvanishing angular momentum density. However, we argue that the qualitative results found in the Georgi-Glashow model are generic and thus a magnetic monopole cannot spin as long as the matter fields feature the usual “monopole” asymptotic behavior independently of the dynamics of the model. A study of the properties of the dyons and magnetic dipoles in some generalizations of the Georgi-Glashow model supplemented with higher order Skyrme-like terms in the gauge curvature and Higgs fields is given quantitatively.

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I. INTRODUCTION AND MOTIVATION

The existence of soliton solutions is an interesting feature of some nonlinear field theory models. Solitons behave like particles, the fields being localized in smooth concentrations of energy density. Moreover, some static solitons are topologically stable (see [1] for a review of these aspects). The systematic study of solitons can be traced back at least to the work of Skyrme [2]. The Skyrme model contains scalar fields (subject to a constraint) only, being proposed as an effective theory for nucleons. In $3 + 1$ dimensional Minkowski spacetime (the case of interest in this work), solitons exist also in models featuring non-Abelian gauge fields.¹ The most prominent such configurations are the 't Hooft-Polyakov magnetic monopoles in the Georgi-Glashow (GG) model [5,6]. Both Skyrmions and monopoles [5,6] are static topologically stable solitons. There exists also a tower of (excited) monopoles [7], generalizing the monopoles of the GG model, each arising from the dimensional descent of the p th member of the Yang-Mills hierarchy [8] on $\mathbb{R}^3 \times S^{4p-3}$ to \mathbb{R}^3 .

¹It is worth noting that no regular particlelike solutions of the pure Yang-Mills (YM) equations exist in Minkowski spacetime [3]. Physically this can be understood as a consequence of the repulsive nature of the YM vector fields. (When scalar fields are added, the existence of solitons become possible due to the balance of the YM repulsive force and the attractive force of the scalars.) However, the no-go results in [3] are circumvented when including gravity effects, as shown by the Bartnik and McKinnon (BK) family of solutions with YM matter fields only [4].

Contrasting with these are the sphalerons in the electro-weak sector of the standard model [9], which have finite energy but are not topologically stable. In the present work, this distinction between topologically stable and sphaleron-like solutions will play a central role.

The 't Hooft-Polyakov monopoles and their Julia-Zee dyonic generalizations [10] are static and spherically symmetric. Then it is natural to wonder whether they possess axially symmetric generalizations with nonzero angular momentum. This question has recently been addressed in the literature, finding a rather unexpected answer. First, it turns out [11,12] that the total angular momentum of the solitons endowed with a net magnetic (topological) charge vanishes (despite the fact that their angular momentum density can be nonzero). Second, the angular momentum of a spinning magnetic dipole (or, more generally, of a configuration with a vanishing net magnetic charge) is nonzero [13] being proportional with the total electric charge.

The pivotal mechanism leading to this conclusion is the fact that the angular momentum density becomes a total divergence by virtue of the electric component of the Euler-Lagrange equations, and, the resulting surface integral then yields vanishing global angular momentum when magnetic monopole boundary values are applied.

The central physical question which we propose to address in this work concerns the generality of these results, namely that solitons with nonvanishing magnetic monopole charge do not spin, irrespective of the specific

dynamical model in question. In practice, the boundary values mentioned above result from the presence of standard, i.e. *quadratic* “kinetic” terms of the Higgs and Yang-Mills fields in the Lagrangian. However, modifications of standard field theoretic models can result in new nonperturbative effects. Examples of this in $3 + 1$ dimensions are, for example, (i) gravity coupled to nonlinear electrodynamics [14], which results in a regular black hole, (ii) the existence of glueballs in a non-Abelian Born-Infeld theory [15], and (iii) the stabilizing of black holes in Einstein-Yang-Mills theory augmented with higher order Yang-Mills curvature terms [16]. Moreover, in higher dimensions, the systematic introduction of higher order YM curvatures and their dimensional descendants results in the construction of instantons and monopoles in all (possible) dimensions [7].

Seeing that more general models than the standard ones with quadratic kinetic terms only may result in qualitatively different solutions, one cannot *a priori* exclude the possibility that a more general choice of the Yang-Mills-Higgs (YMH) Lagrangian might invalidate the results in [11,12] on the general connection between the angular momentum and the electric and magnetic charges. Moreover, by considering generic extensions of YMH systems, we hope to shed some light on the problem of spinning solitons with non-Abelian fields in general.

In what follows, we shall address this question by considering several different generalizations of the GG model and computing the corresponding angular momentum for a given choice of boundary conditions at infinity, reflecting the presence or not of a magnetic charge. Quantitative results are also shown for a specific choice of the YMH Lagrangian. Within the terminology in this work, the configurations with a model with net magnetic and electric charges are called *generalized dyons*. We shall also consider composite configurations with an electric charge only, corresponding to *generalized dipoles*.

This paper is structured as follows. A discussion of the general aspects of the problem of spinning solutions with YMH fields is given in Sec. II, where we exhibit the general framework of the problem. A new result there is a proof that the total angular momentum of a generic YMH system, evaluated by employing the canonical energy momentum tensor, can be expressed as a surface integral at infinity. We continue with Sec. III, where we discuss the basic properties of several generalizations of the GG model. Most of the work in this paper is done for a model where the usual the GG Lagrangian is supplemented with higher order curvature terms of the gauge field and Skyrme-like terms of the gauged Higgs. Our choice of the supplementary part of the Lagrangian has a natural justification, since it can be viewed as the second term ($p = 2$) in a hierarchy of models descended from $4p$ dimensional YM systems [8], the first one of which, $p = 1$, is the usual GG model in the Bogomol’nyi-Prasad-Sommerfield (BPS) limit. Apart from

that, we shall consider also a YMH model with the quadratic YM Lagrangian replaced by a non-Abelian Born-Infeld term, and a YMH model recently introduced in [17], featuring an extra Chern-Simons-like term providing a supplementary interaction between the gauge and Higgs sectors of the theory. The main result of this work is contained in Sec. IV, where we argue that the general connection between the angular momentum and the electric and magnetic charges found in [11,12] for the GG model still holds as long as asymptotically the Higgs field approaches a constant (nonzero) value and the gauge derivative of the Higgs field vanishes. In particular, a magnetic monopole does not possess generalizations with a nonzero total angular momentum. We continue in Sec. V with a quantitative study of the generalized dyons and generalized dipoles in some limits of the $(p=1)+(p=2)$ YMH model. By solving numerically the corresponding field equations, we study how these known solutions of the GG model are affected by the presence of higher derivative YMH terms in the Lagrangian. The last section of this work is devoted to a summary and a discussion of our results. Several possible extensions of the results in Sec. IV are also mentioned there. In the Appendix, we present a peculiar electrically neutral solution of a system consisting of the usual F^2 YM term, plus the pure $p = 2$ YMH system, which might signal the circumvention of the no-angular momentum conjecture. However, it is shown there that the “dyon” of this system must exhibit “magnetic monopole” boundary values and hence the ban on angular momentum for topologically stable YMH solitons cannot be circumvented in this way, supporting our explanation in Sec. VI.

II. THE FORMALISM

A. The conventions and notations

In this work we shall ignore the backreaction of the matter fields on the geometry and consider a fixed four-dimensional Minkowski spacetime background,² with a line element $ds^2 = dt^2 - (dx^2 + dy^2 + dz^2)$, where t is the time coordinate and x, y, z are the usual cartesian coordinates. The same line element expressed in spherical coordinates reads $ds^2 = dt^2 - dr^2 - r^2(d\theta^2 + \sin^2\theta d\varphi^2)$, where $0 \leq r < \infty$ is the radial coordinate and θ, φ are the spherical coordinates on S^2 , with the usual range. For completeness, we also give the corresponding expression in cylindrical coordinates, $ds^2 = dt^2 - d\rho^2 - \rho^2 d\varphi^2 - dz^2$, where $0 \leq \rho < \infty$. Note that in all relations, the greek indices (like μ, ν) are running from 0 to 3 (with $x^0 = t$).

The gauge potential A_μ and Higgs field Φ are denoted as

$$A_\mu = -\frac{i}{2} A_\mu^a \tau_a, \quad \Phi = -\frac{i}{2} \Phi^a \tau_a, \quad (1)$$

²See, however, the remarks in the last section on the gravity effects.

with τ_a the Pauli matrices ($a = 1, 2, 3$). The resulting anti-Hermitian curvature and covariant derivative are then

$$\begin{aligned} F_{\mu\nu} &= \partial_\mu A_\nu - \partial_\nu A_\mu + [A_\mu, A_\nu], \\ D_\mu \Phi &= \partial_\mu \Phi + [A_\mu, \Phi]. \end{aligned} \quad (2)$$

Also, for the purposes of this work we found it convenient to introduce the notation

$$\begin{aligned} \mathcal{F}_{\mu\nu\rho\sigma} &= \{F_{\mu[\nu}, F_{\rho\sigma]}\} = F_{\mu\nu\rho\sigma}, \\ \mathcal{F}_{\mu\nu\rho} &= \{F_{[\mu\nu}, D_{\rho]}\Phi\}, \\ \mathcal{F}_{\mu\nu} &= \{S, F_{\mu\nu}\} + [D_\mu\Phi, D_\nu\Phi], \\ \mathcal{F}_\mu &= \{S, D_\mu\Phi\}, \\ \mathcal{F} &= S^2, \quad \text{with } S \stackrel{\text{def}}{=} -(\eta^2 \mathbb{1} + \Phi^2), \end{aligned} \quad (3)$$

where we have used $[ijk]$ to denote cyclic symmetry in the indices i, j, k and $\{A, B\}$ denotes an anticommutator. The generalizations of the GG Lagrangian we shall consider in what follows are built in terms of Eq. (3) above.

B. The issue of axial symmetry

The physical configurations we are interested in have no time dependence and are axially symmetric (i.e. they remain invariant under a rotation around the z axis). This implies the existence of two Killing vectors of the problem $\partial/\partial t$ and $\partial/\partial\varphi$. For an Abelian gauge field this implies directly that the gauge potential has no dependence on t and φ . However, the issue of interplay between spacetime and gauge symmetries is rather subtle in the presence of non-Abelian matter fields.³ In this case, the symmetry of the gauge field under a spacetime symmetry (as characterized by a given Killing vector) means that the action of an isometry can be compensated by a suitable gauge transformation [19,20]. For the Killing vector $\partial/\partial t$, the natural choice is to choose a gauge with $\partial A/\partial t = \partial\Phi/\partial t = 0$. However, a rotation around the z axis can be compensated by a gauge rotation, $\partial_\varphi A_\mu = D_\mu\psi$ (with ψ an element of the algebra), and therefore

$$F_{\mu\varphi} = D_\mu W, \quad D_\varphi \Phi = [W, \Phi], \quad (4)$$

where

$$W = A_\varphi + \psi. \quad (5)$$

These relations (which are independent of the choice of a specific YMH Lagrangian) allow us in what follows to express the angular momentum density as a flux integral at infinity.

³A nice discussion of the relationship between conservation laws, spacetime symmetries, and gauge symmetries can be found in Ref. [18].

C. The energy-momentum tensor and general relations

The Lagrangian \mathcal{L} of the model (which is not specified at this stage) is a function of the matter fields $\Psi = (A_\mu, \Phi)$. However, the requirement that the equations of motion are of second order plus the gauge covariance implies that \mathcal{L} depends only on $F_{\mu\nu}$, $D_\mu\Phi$, and Φ .

We start by defining the generalized momentum densities

$$\Pi^{\mu\nu} = \frac{\partial\mathcal{L}}{\partial A_{\nu,\mu}}, \quad \Pi^\mu = \frac{\partial\mathcal{L}}{\partial\Phi_{,\mu}}, \quad (6)$$

such that the Euler-Lagrange equations $\delta_\Psi\mathcal{L} = 0$ can be written as⁴

$$D_\mu\Pi^{\mu\nu} - [\phi, \Pi^\nu] = 0, \quad \Pi^\mu_{,\mu} = \frac{\partial\mathcal{L}}{\partial\Phi}. \quad (7)$$

As usual, the invariance of the action under four translations, $x^\mu \rightarrow x^\mu + a^\mu$ yields the canonical energy-momentum tensor [21]

$$\begin{aligned} 2T^\beta_\alpha &= \text{Tr}\left(A_{\mu,\alpha}\frac{\partial\mathcal{L}}{\partial A_{\mu,\beta}} + \Phi_{,\alpha}\frac{\partial\mathcal{L}}{\partial\Phi_{,\beta}}\right) - \delta^\beta_\alpha\mathcal{L} \\ &= \text{Tr}(A_{\mu,\alpha}\Pi^{\beta\mu} + \Phi_{,\alpha}\Pi^\beta) - \delta^\beta_\alpha\mathcal{L}, \end{aligned} \quad (8)$$

which is conserved, $T^\nu_{,\nu} = 0$. Following the usual prescription, this expression can be made gauge invariant by adding to it a total divergence [22], such that

$$2T^\beta_\alpha = \text{Tr}(F_{\alpha\mu}\Pi^{\beta\mu} + D_\alpha\Phi\Pi^\beta) - \delta^\beta_\alpha\mathcal{L}. \quad (9)$$

At this point is worth recalling that the canonical energy momentum tensor $T_{\mu\nu}$ suffers from a number of well-known problems; for example it is not explicitly symmetric in the indices μ, ν [this holds also for Eq. (9)]. As usual, an energy-momentum tensor which is directly symmetric and gauge invariant is found by introducing the spacetime metric $g_{\mu\nu}$ into the action and assuming it to be arbitrary; then the energy-momentum tensor is obtained by differentiating the density of the action with respect to the metric:

$$T_{\alpha\beta} = \frac{2}{\sqrt{-g}} \frac{\delta\sqrt{-g}\mathcal{L}}{\delta g^{\alpha\beta}}. \quad (10)$$

One can easily show that the energy-momentum tensor and the canonical one are identical for the case of a GG model. However, we have verified that the expression of $T_{\alpha\beta}$ obtained via the definition Eq. (9) and via Eq. (10) also coincide for the specific models Eqs. (21), (29), and (31) below.

⁴Note that these relations are given in a Cartesian coordinate system.

The advantage of using the definition Eq. (9) is that it leads to an expression of the angular momentum density, T_φ^t , as a total divergence, independent on the choice of \mathcal{L} . In proving that, we make use of the general relations Eq. (4) together with the generalized YM equations Eq. (7). After replacing in Eq. (9), one finds the following general expression⁵ of T_φ^t :

$$2T_\varphi^t = \text{Tr} \left(\frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^\mu} (\sqrt{-g} W \Pi^{\mu t}) \right). \quad (11)$$

As a result, the total angular momentum can be expressed as a flux integral⁶

$$J = \int d^3x \sqrt{-g} T_\varphi^t = \frac{1}{2} \oint_\infty d\Sigma_k \text{Tr} (\sqrt{-g} W \Pi^{kt}). \quad (12)$$

Therefore we conclude that in a general YMH theory only the large- r asymptotic structure of the fields is relevant for the issue of angular momentum of regular configurations. (Note that this relation has been proven in [11] for a GG model.)

For the purposes of this work, it is also useful to define the *electric* part of the energy density as

$$2\mathcal{E} = \text{Tr} (F_{at} \Pi^{at} + D_t \Phi \Pi^t). \quad (13)$$

Similar to the angular momentum density, one can show that Eq. (13) can also be written as total divergence

$$2\mathcal{E} = \text{Tr} \left(\frac{1}{\sqrt{-g}} \frac{\partial}{\partial x^\mu} (\sqrt{-g} A_t \Pi^{\mu t}) \right). \quad (14)$$

Then the total *electric mass* can also be expressed as a surface integral

$$E_e = \int d^3x \sqrt{-g} \mathcal{E} = \frac{1}{2} \oint_\infty d\Sigma_k \text{Tr} (\sqrt{-g} A_t \Pi^{kt}). \quad (15)$$

The electric charge can also be expressed as a surface integral, a natural definition of it being

$$Q_e = \frac{1}{4\pi} \oint_\infty dS_k \text{Tr} (\hat{\Phi} \Pi^{kt}), \quad (16)$$

for all YMH models to be discussed below (where we define as usual $\hat{\Phi} = \Phi/|\Phi|$).

The definition of the magnetic charge is dependent of the particular model, and these magnetic monopole charge densities will be defined for each model in turn below.

⁵Here we change to a spherical (or cylindrical) coordinate system, such that $\mu = r, \theta$ (or $\mu = \rho, z$, respectively), g being the determinant of the metric tensor.

⁶Note that Eq. (12) holds also for higher gauge groups, since it relies on the general relation Eq. (4).

III. THE SPECIFIC MODELS

A. The Georgi-Glashow model on $\mathbb{R}^{3,1}$: The $p = 1$ model

The ‘‘canonical’’ model of a YMH theory is the GG one, with a Lagrangian

$$\begin{aligned} \mathcal{L}^{(1)} = & -\frac{1}{4} \text{Tr} (F_{\mu\nu} F^{\mu\nu}) - \frac{1}{2} \text{Tr} (D_\mu \Phi D^\mu \Phi) \\ & + \frac{\lambda}{4} (|\Phi^a|^2 - \eta^2)^2, \end{aligned} \quad (17)$$

the last term being the symmetry-breaking Higgs self-iteration potential. η is the scale-breaking Higgs vacuum expectation value with dimensions of L^{-1} . In what follows, we will consider the Bogomol’nyi-Prasad-Sommerfield (BPS) limit $\lambda = 0$ case exclusively. This choice is made partly for simplifying the analysis, but also because the dimensional reduction of the (usual) YM system on $\mathbb{R}^3 \times S^1$ yields the $p = 1$ YMH system in the BPS limit.⁷

For the Lagrangian Eq. (17),

$$\Pi^{\mu\nu} = F^{\mu\nu}, \quad \Pi^\mu = D^\mu \Phi. \quad (18)$$

The stress tensor of this model, as read from Eq. (9) is [here we give the (more transparent) expression in terms of $F_{\mu\nu}$ and $D_\mu \Phi$]

$$\begin{aligned} 2T_{\mu\nu}^{(1)} = & \text{Tr} \left(F_{\mu\tau} F_\nu{}^\tau - \frac{1}{4} g_{\mu\nu} F_{\tau\lambda} F^{\tau\lambda} \right) \\ & - \text{Tr} \left(D_\mu \Phi D_\nu \Phi - \frac{1}{2} g_{\mu\nu} D_\tau \Phi D^\tau \Phi \right). \end{aligned} \quad (19)$$

The GG model has a variety of interesting features which have been extensively studied in the literature over the last 40 years. Here we mention only that in the static, purely magnetic limit (i.e. $A_t = 0$), the Hamiltonian of the model (which in that case coincides with the Lagrangian) is bounded from below by the magnetic charge density

$$Q^{(1)} = \frac{1}{4\pi} \varepsilon_{ijk} \text{Tr} (F_{ij} D_k \Phi) \stackrel{\text{def}}{=} \mathbf{\nabla} \cdot \mathbf{\Omega}^{(1)},$$

$$\text{with } \mathbf{\Omega}^{(1)} = \frac{1}{4\pi} \varepsilon_{ijk} \text{Tr} (\Phi F_{ij}), \quad (20)$$

which is the dimensional descendant of the second Chern-Pontryagin density. The total magnetic charge Q_m is the integral of $Q^{(1)}$.

B. The $(p = 1) + (p = 2)$ model

In our choice of this YMH model, we start from the simple observation that the GG model Eq. (17) can be

⁷This is true also of the $p = 2$ YMH system resulting from the dimensional reduction of the $p = 2$ YM system on $\mathbb{R}^3 \times S^1$, though in that case the BPS limit cannot be saturated.

obtained as descending from the pure F^2 YM Lagrangian on $R^3 \times S^1$, the components of the YM potential along the S^1 direction corresponding to the Higgs fields. This observation has led in [8] to the construction of a hierarchy of $p \geq 1$ $SO(3)$ Higgs models in $D = 3$ space dimensions, obtained by dimensional descent over S^{4p-3} , of the p th member of the Yang-Mills hierarchy. The $p = 1$ member here is nothing else than the BPS limit of the GG model. For any $p \geq 2$, the Lagrangian $\mathcal{L}^{(p)}$ is a Skyrme-like gauged Higgs system, in the sense that it consists of $2p$ th powers of the gauge curvature $F(2)$ and covariant derivative $D\Phi$, suitably antisymmetrized so that only the squares of “velocity” fields appear. Moreover, for each $p \geq 1$ YMH model, one can define a topological charge density which is the descendant of the $2p$ th Chern-Pontryagin density. In each case there exists a Bogolmol’nyi bound, but this bound can be saturated only in the $p = 1$ case.

In the present work, we restrict our attention to the system described by the Lagrangian of the $(p = 1) + (p = 2)$ model, which consists of the sum of the GG model in the BPS limit (i.e. $p = 1$) plus a “correction part,” which is inspired by the Lagrangian of the $p = 2$ term in the general YMH hierarchy introduced in [8]

$$\mathcal{L} = \mathcal{L}^{(1)} + \mathcal{L}^{(2)}. \quad (21)$$

It is clear that restricting to “corrections” of the $p = 2$ YMH is sufficient to describe the effect of such correction qualitatively. The $p = 2$ YMH model is described in detail in the next subsection.

1. The $p = 2$ YMH system on $\mathbb{R}^{3,1}$

The Lagrangian of the general $p = 2$ model can be written as a sum of five different terms

$$\frac{1}{20}\mathcal{L}^{(2)} = \sum_{a=0}^4 \lambda_a \mathcal{L}^{(2,a)}, \quad (22)$$

with

$$\begin{aligned} \frac{1}{20}T_{\mu\nu}^{(2)} = & -\lambda_0 \text{Tr} \left(F_{\mu\tau\lambda\sigma} F_{\nu}{}^{\tau\lambda\sigma} - \frac{1}{8} g_{\mu\nu} F_{\kappa\tau\lambda\sigma} F^{\kappa\tau\lambda\sigma} \right) + 3\lambda_1 \text{Tr} \left(\{F_{[\mu\tau}, D_{\lambda]} \Phi\} \{F_{[\nu}{}^{\tau}, D^{\lambda]} \Phi\} - \frac{1}{6} g_{\mu\nu} \{F_{[\mu\nu}, D_{\rho]} \Phi\} \{F^{[\mu\nu}, D^{\rho]} \Phi\} \right) \\ & + 2 \cdot 6\lambda_2 \text{Tr} \left((\{S, F_{\mu\tau}\} + [D_{\mu} \Phi, D_{\tau} \Phi]) (\{S, F_{\nu}{}^{\tau}\} + [D_{\nu} \Phi, D^{\tau} \Phi]) - \frac{1}{4} g_{\mu\nu} (\{S, F_{\tau\lambda}\} + [D_{\tau} \Phi, D_{\lambda} \Phi]) (\{S, F^{\tau\lambda}\} + [D^{\tau} \Phi, D^{\lambda} \Phi]) \right) \\ & - 9\lambda_3 \text{Tr} \left(\{S, D_{\mu} \Phi\} \{S, D_{\nu} \Phi\} - \frac{1}{2} g_{\mu\nu} \{S, D_{\lambda} \Phi\} \{S, D^{\lambda} \Phi\} \right) - 54\lambda_4 \text{Tr} \left(0 - \frac{1}{2} g_{\mu\nu} S^4 \right). \end{aligned} \quad (26)$$

⁸In what follows, we take $\lambda_a \geq 0$ as implied by requiring a strictly positive energy density.

⁹As discussed in [16], a $\mathcal{F}_{\mu\nu\rho\sigma} \mathcal{F}^{\mu\nu\rho\sigma}$ term stabilizes the Einstein–Yang–Mills hairy black holes, leading to solutions with very different properties as compared to the standard ones [23]. Moreover, this term leads to new qualitative features in inflationary models with non-Abelian gauge fields [24].

$$\begin{aligned} \mathcal{L}^{(2,0)} &= -\frac{1}{4} \mathcal{F}_{\mu\nu\rho\sigma} \mathcal{F}^{\mu\nu\rho\sigma}, & \mathcal{L}^{(2,1)} &= \mathcal{F}_{\mu\nu\rho} \mathcal{F}^{\mu\nu\rho}, \\ \mathcal{L}^{(2,2)} &= 6\mathcal{F}_{\mu\nu} \mathcal{F}^{\mu\nu}, & \mathcal{L}^{(2,3)} &= -9\mathcal{F}_{\mu} \mathcal{F}^{\mu}, \\ \mathcal{L}^{(2,4)} &= -54\mathcal{F}^2, \end{aligned} \quad (23)$$

where we have used the symbolic notation stated in Eq. (3).

It is clear that the Lagrangian Eq. (22), is not renormalizable and can be analyzed only at the classical level. It can be seen as an effective action of a renormalizable theory at low energy. In this respect it is analogous to the Skyrme model [2] versus the quadratic $O(4)$ sigma model, and in this respect can be viewed as the Yang–Mills–Higgs analogue of that sigma model.

The Lagrangian of the original $p = 2$ YMH model (obtained by dimensional descent from a pure F^4 Yang–Mills model in $D = 8$ dimensions) is found by taking $\lambda_0 = 0, \lambda_1 = \lambda_2 = \lambda_3 = \lambda_4 = 1$. However, it is of interest to treat λ_a as arbitrary positive parameters⁸ and to work with the generic model Eq. (22). Moreover, the term $\mathcal{F}_{\mu\nu\rho\sigma} \mathcal{F}^{\mu\nu\rho\sigma}$ has been added by hand in Eq. (22) since it does not occur in the original $p = 2$ YMH model. Note that in the purely magnetic limit, this term is absent, and therefore it is interesting to see what effect its inclusion has in the construction of the electrically charged solutions.⁹

The expressions of the generalized momentum densities for each of the terms in Eq. (22), is

$$\begin{aligned} \Pi_{(2,0)}^{\mu\nu} &= -3! \{ \mathcal{F}^{\mu\nu\rho\sigma}, F_{\rho\sigma} \}, & \Pi_{(2,1)}^{\mu\nu} &= 2 \cdot 3! \{ \mathcal{F}^{\mu\nu\rho}, D_{\rho} \Phi \}, \\ \Pi_{(2,2)}^{\mu\nu} &= -4! \{ S, \mathcal{F}^{\mu\nu} \}, \end{aligned} \quad (24)$$

and

$$\begin{aligned} \Pi_{(2,1)}^{\mu} &= 3! \{ \mathcal{F}^{\mu\rho\sigma}, F_{\rho\sigma} \}, & \Pi_{(2,2)}^{\mu} &= -4! [\mathcal{F}^{\mu\rho}, D_{\rho} \Phi], \\ \Pi_{(2,3)}^{\mu} &= -18 \{ S, \mathcal{F}^{\mu} \}. \end{aligned} \quad (25)$$

The corresponding symmetric and gauge-invariant energy-momentum tensor reads (here we give the final expression used in practice, in terms of $F_{\mu\nu}$ and $D_{\mu} \Phi$)

It may be interesting to display the (magnetic) topological charge density of the $p = 2$ YMH model, alluded to in the previous subsection,

$$\begin{aligned} \varrho^{(2)} &\simeq \varepsilon_{ijk} \text{Tr} (3\eta^4 F_{ij} D_k \Phi + \eta^2 [3F_{ij} (\Phi^2 D_k \Phi + D_k \Phi \Phi^2) - 2D_i \Phi D_j \Phi D_k \Phi]) \\ &\quad + [F_{ij} (\Phi^4 D_k \Phi + D_k \Phi \Phi^4 + \Phi^2 D_k \Phi \Phi^2) - 2\Phi^2 D_i \Phi D_j \Phi D_k \Phi] \\ &\stackrel{\text{def}}{=} \nabla \cdot \Omega^{(2)}. \end{aligned} \quad (27)$$

This pertains to the system Eq. (22), (with $\lambda_0 = 0$, $\lambda_1 = \lambda_2 = \lambda_3 = \lambda_4 = 1$) on \mathbb{R}^3 and descends from the fourth Chern-Pontryagin density in 8 dimensions. It is manifestly a total divergence

$$\Omega_k^{(2)} \simeq \varepsilon_{ijk} \text{Tr} \left[3\eta^4 \Phi F_{ij} + 2\eta^2 (\Phi^3 F_{ij} - \Phi D_i \Phi D_j \Phi) + \frac{1}{5} [3\Phi^5 F_{ij} - 2(2\Phi^3 D_i \Phi D_j \Phi - \Phi^2 D_i \Phi \Phi D_j \Phi)] \right]. \quad (28)$$

It is interesting to note that the surface integrals of Eqs. (28) and (20), are equal, up to a numerical multiple. This is because the terms in Eq. (28), not featuring the curvature F_{ij} decay too fast to contribute, by virtue of Higgs asymptotics. This density was employed in [25,26].

C. The non-Abelian Born-Infeld–Higgs model

Another interesting possibility is to consider a Born-Infeld (BI) Lagrangian for the gauge fields. This modification of the standard YM quadratic Lagrangian is suggested by the superstring theory [27,28] leading to a variety of interesting features. For example, the no-go results in [3] forbidding the existence of pure YM solitons are circumvented in this case, since a non-Abelian BI theory possesses particlelike solutions even in the absence of a Higgs field [15]. The spherically symmetric monopoles and dyons of this model have been studied in [29,30].

The Lagrangian of the non-Abelian BI–Higgs theory reads

$$\begin{aligned} \mathcal{L}^{(\text{BIH})} &= \beta^2 (1 - \sqrt{1 + U}) - \frac{1}{2} \text{Tr} (D_\mu \Phi D^\mu \Phi), \\ \text{with } U &= \frac{1}{2\beta^2} \text{Tr} (F_{\mu\nu} F^{\mu\nu}) - \frac{1}{\beta^4} \text{Tr} (F_{\mu\nu\rho\sigma} F^{\mu\nu\rho\sigma}), \end{aligned} \quad (29)$$

where Tr represents the usual trace on $SO(3)$ indices. [Note that the definition of the BI theory for a non-Abelian gauge group is not unique and several alternatives have been discussed in the literature. Apart from the one used above, another possibility of interest (which will shall not consider here) is to take a symmetric trace operation [27], but so far the explicit Lagrangian with such trace is known only as perturbative series.]

The corresponding expressions of the generalized momentum densities read

$$\Pi^{\mu\nu} = -\frac{1}{\sqrt{1+U}} \left(F^{\mu\nu} - \frac{2}{\beta^2} \{ \mathcal{F}^{\mu\nu\rho\sigma}, F_{\rho\sigma} \} \right), \quad \Pi^\mu = D^\mu \Phi. \quad (30)$$

In the absence of a topological lower bound, it is natural to use the magnetic charge definition valid for the GG model, also in this case.

D. Yang-Mills–Higgs model with Chern-Simons–like term

In 2 + 1 dimensions, electric charge and angular momentum result [31–33] from the dynamics of a Chern-Simons term in the Lagrangian. In 3 + 1 dimensions, however, no Chern-Simons density was identified until recently proposed in [7,34]. In a given spacetime dimension, they result from the descent by one step, of the 1-form in the total divergence expression of a Chern-Pontryagin (CP) density, which itself is a dimensional descendant of a CP density in some higher (even) dimension. These new Chern-Simons densities feature both gauge fields and Higgs fields, and are defined in both odd and even spacetime dimensions.

Here, we consider the simplest example in 3 + 1 dimensional spacetime, which was discussed recently in [17]. This CS density is extracted from the dimensionally reduced CP density on $\mathbf{M}^5 \times S^1$. The residual CP density on \mathbf{M}^5 being a total divergence, a Chern-Simons density in 3 + 1 dimensions can be extracted in the usual way. The residual gauge group then is $SO(5)$ and the Higgs field takes its values in the algebra of $SO(6)$.

The Lagrangian density of this specific Yang-Mills–Higgs–Chern-Simons (YMHCS) model reads

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4} \text{Tr} (F_{\mu\nu} F^{\mu\nu}) - \frac{1}{2} \text{Tr} (D_\mu \Phi D^\mu \Phi) \\ &\quad + i\kappa \varepsilon^{\mu\nu\rho\sigma} \text{Tr} (\Phi F_{\mu\nu} F_{\rho\sigma}), \end{aligned} \quad (31)$$

with κ an arbitrary constant. The generalized momentum densities are

$$\Pi^{\mu\nu} = F^{\mu\nu} + i\kappa\epsilon^{\mu\nu\rho\sigma}\text{Tr}(\Phi F_{\rho\sigma}), \quad \Pi^\mu = D^\mu\Phi. \quad (32)$$

Following [17], we shall restrict our study to Yang-Mills fields taking their values in the $SO(3) \times U(1)$ subalgebra of $SO(5)$, with an $SO(3)$ Higgs triplet. In this limit, the Lagrangian Eq. (31) of the model can be written in an equivalent form as

$$\begin{aligned} \mathcal{L} = & -\frac{1}{4}\text{Tr}(F_{\mu\nu}F^{\mu\nu}) - \frac{1}{2}\text{Tr}(D_\mu\Phi D^\mu\Phi) - \frac{1}{4}f_{\mu\nu}f^{\mu\nu} \\ & + i\kappa\epsilon^{\mu\nu\rho\sigma}f_{\mu\nu}\text{Tr}(\Phi F_{\rho\sigma}), \end{aligned} \quad (33)$$

with $f_{\mu\nu} = \partial_\mu a_\nu - \partial_\nu a_\mu$ a $U(1)$ field and $F_{\mu\nu}$, Φ the usual $SO(3)$ gauge fields and Higgs field, respectively. As found in [17], this extra term prevents the existence of (self-dual) solutions saturating a lower bound.

The Lagrangian Eq. (31) corresponds to the flat space-time limit of a YMH-Maxwell model considered in Sec. II of [35] in a more general context. Interestingly, the results in [35] show that when including the gravity effect, the self-gravitating solitonic solutions of Eq. (33) saturate a gravitational version of the Bogomol'nyi bound, with the same magnetic charge definition as in the GG model.

IV. THE ISSUE OF SPINNING SOLUTIONS

A. Axially symmetric YMH fields

In order to evaluate the general relation Eq. (12) and to construct generalized dyons and dipoles, we need to specify a YMH ansatz. Employing spherical coordinates, we use the following parametrization of the axially symmetric YMH fields:

$$\begin{aligned} A_\mu dx^\mu = & \left(H_5 \frac{i\tau_r^{(n)}}{2} + H_6 \frac{i\tau_\theta^{(n)}}{2} \right) dt \\ & + \left[\frac{H_1}{r} dr + (1 - H_2) d\theta \right] \frac{i\tau_\varphi^{(n)}}{2} \\ & - n \sin\theta \left(H_3 \frac{i\tau_r^{(n)}}{2} + H_4 \frac{i\tau_\theta^{(n)}}{2} \right) d\varphi, \end{aligned} \quad (34)$$

$$\Phi = \eta \left(\Phi_1 \frac{i\tau_r^{(n)}}{2} + \Phi_2 \frac{i\tau_\theta^{(n)}}{2} \right), \quad (35)$$

in terms of six gauge potentials and two Higgs functions. The only φ -dependent terms in Eq. (34) are the $SU(2)$ matrices $\tau_r^{(n)}$, $\tau_\theta^{(n)}$, and $\tau_\varphi^{(n)}$. These matrices are defined as $\tau_r^{(n)} = \sin\theta(\cos n\varphi\tau_x + \sin n\varphi\tau_y) + \cos\theta\tau_z$, $\tau_\theta^{(n)} = \cos\theta(\cos n\varphi\tau_x + \sin n\varphi\tau_y) - \sin\theta\tau_z$, and $\tau_\varphi^{(n)} = -\sin n\varphi\tau_x + \cos n\varphi\tau_y$ [with $\tau_a = (\tau_x, \tau_y, \tau_z)$ the Pauli matrices]. For

$H_5 = H_6 = 0$, Eq. (34) is essentially the axially symmetric YMH ansatz as introduced by Manton [36], and Rebbi and Rossi [37] when discussing multimono-pole solutions. This particular parametrization of the YM ansatz in terms of H_i is very convenient for numerical studies and is employed in most of the work on axially symmetric YMH systems. Note that Eq. (35) contains, via the matrices $\tau_r^{(n)}$, $\tau_\theta^{(n)}$, and $\tau_\varphi^{(n)}$, an extra integer $n = 1, 2, \dots$ which is the winding number of the solutions.

Let us also mention that the above Ansatz possesses a residual Abelian gauge invariance (see e.g. the discussion in [38]). To fix it, we have to include a gauge-fixing term, the most convenient choice being $r\partial_r H_1 - \partial_\theta H_2 = 0$.

B. The far field asymptotics

For any choice of the Lagrangian, the assumption that the Higgs field approaches asymptotically a constant value $|\Phi| \rightarrow 1$ (which is the effect of a symmetry breaking Higgs potential, whether explicitly included in the action or not), results in the following finite energy condition on the Higgs:

$$\Phi^1 = \cos(m-1)\theta, \quad \Phi^2 = \sin(m-1)\theta, \quad (36)$$

with $m = 1, 2, \dots$ a positive integer. In the next step, we shall assume that the gauge derivative of the Higgs fields vanished asymptotically, $D_\mu\Phi D_\mu\Phi \rightarrow 0$, a condition which then fixes the asymptotic values of the YM potentials. For odd m these are

$$\begin{aligned} H_1 = 0, \quad H_2 = -(m-1), \\ H_3 = \frac{\cos\theta}{\sin\theta}[\cos(m-1)\theta - 1], \quad H_4 = -\frac{\cos\theta}{\sin\theta}\sin(m-1)\theta, \\ H_5 = V_0 \cos(m-1)\theta, \quad H_6 = V_0 \sin(m-1)\theta, \end{aligned} \quad (37)$$

and for even m

$$\begin{aligned} H_1 = 0, \quad H_2 = -(m-1), \\ H_3 = \frac{1}{\sin\theta}[\cos(m-1)\theta - \cos\theta], \quad H_4 = -\frac{\sin(m-1)\theta}{\sin\theta}, \\ H_5 = V_0 \cos(m-1)\theta, \quad H_6 = V_0 \sin(m-1)\theta, \end{aligned} \quad (38)$$

with V_0 a constant. (A derivation of these boundary conditions can be found in [12,39,40].) Thus, for $m = 2k$ ($k = 1, 2, \dots$), the ground state of the model corresponds to a gauge transformed trivial solution and the magnetic charge vanishes. The situation is different for odd values $m = 2k + 1 = 1, 3, \dots$, the ground state in this case corresponding to a charge n multimono-pole.

To evaluate Eq. (12), we need both W , which for the axially symmetric configuration at hand is

$$W = -n(\cos\theta + \sin\theta H_3) \frac{i\tau_r^{(n)}}{2} + n \sin\theta H_4 \frac{i\tau_\theta^{(n)}}{2}, \quad (39)$$

and, the asymptotic expression of the generalized momentum Π^{rt} . At this point, we remark that, for all models in this work, the YM Lagrangian effectively includes the quadratic term $F_{\mu\nu}^2$, namely the $p = 1$ YM term. Now the additional curvature YM terms are all higher order in the curvature 2-form, so they decay faster than the usual $p = 1$, and hence their contributions to the Eq. (12), integral will vanish. Then the following expression holds for large r : $\Pi^{rt} = F^{rt} + (\text{subleading terms})$, and the electric potentials H_5 and H_6 have long-range, Coulomb-like tails,

$$\begin{aligned} H_5 &= \cos(m-1)\theta \left(V_0 - \frac{Q_e}{r} \right) + \dots, \\ H_6 &= \sin(m-1)\theta \left(V_0 - \frac{Q_e}{r} \right) + \dots, \end{aligned} \quad (40)$$

with Q_e corresponding to the electric charge as computed from Eq. (16).

We close this part by noticing that the *electric mass* Eq. (15) can always be written as a product between the “electrostatic potential” V_0 and the electric charge Q_e ,

$$E_e = \frac{1}{2} V_0 Q_e. \quad (41)$$

Assuming that the *electric* part of the energy density \mathcal{E} , as given by Eq. (13), is a positive quantity, it follows that a configuration with $V_0 = 0$ or $Q_e = 0$ does not possess an electric field $F_{kt} \equiv 0$.

Another important feature is that the “electrostatic potential” V_0 (as fixed by the asymptotics of the functions H_5, H_6) is always bounded from above by the vev of the Higgs field (i.e. $V_0 < 1$ for the conventions here). Since the $p = 1$ term would dominate as $r \rightarrow \infty$, a value $V_0 > 1$ implies a far field oscillatory behavior of some YM potentials, a feature which is not compatible with finite energy requirements (this can be seen explicitly in the Bogomol’nyi-Prasad-Sommerfield exact solution [41]; see also [42] for the axially symmetric case).

C. The total angular momentum

With these relations at hand, the evaluation of the general relation Eq. (12) is straightforward. For the parametrization Eq. (34), the asymptotic expression of W which enters the general relation Eq. (4) is

$$W = -n \cos\theta \left(\cos(m-1)\theta \frac{i\tau_r^{(n)}}{2} + \sin(m-1)\theta \frac{i\tau_\theta^{(n)}}{2} \right),$$

for odd m and

$$W = -n \left(\cos(m-1)\theta \frac{i\tau_r^{(n)}}{2} + \sin(m-1)\theta \frac{i\tau_\theta^{(n)}}{2} \right),$$

for even m . Also, for any m , the asymptotic expression of the generalized momentum Π^{rt} is

$$\Pi^{rt} = \left(\cos(m-1)\theta \frac{i\tau_r^{(n)}}{2} + \sin(m-1)\theta \frac{i\tau_\theta^{(n)}}{2} \right) \frac{Q}{r^2} + \dots \quad (42)$$

Then, after replacing Π^{rt} in Eq. (12), one finds that for both ($p = 1$) + ($p = 2$) models and the BI-Higgs models the following relations hold:

$$\begin{aligned} J &= 0 \quad \text{for } m = 1, 3, \dots \quad (Q_M \neq 0), \\ J &= nQ_e \quad \text{for } m = 2, 4, \dots \quad (Q_M = 0). \end{aligned} \quad (43)$$

This is the relation derived before for the pure GG case, connecting the quantization of charge and angular momentum. However, we see that this is a more general result, which holds as long as the YMH fields share the asymptotics of the GG model (while the precise features in the bulk can be very different). Therefore we conclude that a magnetic monopole does not spin while a magnetic dipole cannot rotate unless it is endowed with a net electric charge.

The situation is more involved for the model Eq. (33) featuring a Chern-Simons-like term, since the angular momentum density has a supplementary part originating from this extra term. To evaluate this supplementary contribution, we write the usual axially symmetric ansatz for the U(1) field 1-form

$$a = a_\phi d\phi + a_t dt. \quad (44)$$

The boundary condition by a_μ as $r \rightarrow \infty$ are $a_\phi \rightarrow 0$ and $a_t \rightarrow v_0$ (with v_0 an arbitrary constant). Moreover, one can easily see that the asymptotic behavior of the YMH field remains the same as in the GG model. Then a straightforward computation shows that the general relations Eq. (43) still holds for this model. Thus we conclude that the supplementary YMHCS-like term in Eq. (33) does not endow a monopole with a nonzero angular momentum.

However, one should remark that the charge-parity violating term in Eq. (33), is not strictly speaking a Chern-Simons term, unlike the corresponding term in Eq. (31). Thus, any statements concerning the vanishing of the angular momentum is not *a priori* valid for the general Chern-Simons theory Eq. (31). It is conceivable that in that case the angular momentum may not vanish. Any progress in this direction requires first an investigation of the issue of spinning monopoles for higher gauge groups and for different representations of the Higgs field, a task which has not been yet considered in the literature.

V. GENERALIZED DYONS AND MAGNETIC DIPOLES IN $(p = 1) + (p = 2)$ YMH MODEL: NUMERICAL RESULTS

Although we have seen that all solutions share the same relation between the angular momentum and electric and magnetic charges as in the GG model, it remains an interesting question to see how a more general YMH Lagrangian *quantitatively* affects the properties of the known axially symmetric configurations.

To answer this question, we have considered the $(p = 1) + (p = 2)$ YMH model and constructed generalizations of the simplest solutions corresponding to dyons and magnetic dipoles. To our knowledge, this problem has not been addressed before in the literature, only purely

magnetic solutions being considered so far. The monopole solutions of the $(p = 1) + (p = 2)$ model have been discussed in [25,26], with a rather complicated picture being revealed. For example, this model does not support self-dual solutions [43]. Moreover, depending on the values of the dimensionless coupling parameters, the generalized monopoles exhibit both attractive and repulsive phases.

In this section we extend the results of [25] in several different directions, by including an electric component in the YM connection.

A. The boundary conditions

The boundary conditions at infinity for dyons are found by taking $m = 1$ in the general asymptotics Eqs. (36), (37)

$$H_1|_{r=\infty} = H_2|_{r=\infty} = H_3|_{r=\infty} = H_4|_{r=\infty} = H_6|_{r=\infty} = \Phi_2 = 0, \quad H_5|_{r=\infty} = V_0, \Phi_1 = 1, \quad (45)$$

with $V_0 < 1$ a constant corresponding to the electric potential. The corresponding boundary conditions at the origin $r = 0$ are

$$H_1|_{r=0} = H_3|_{r=0} = H_5|_{r=0} = H_6|_{r=0} = \Phi_1|_{r=0} = \Phi_2|_{r=0} = 0, \quad H_2|_{r=0} = H_4|_{r=0} = 1. \quad (46)$$

On the symmetry axis, the dyons satisfy the boundary conditions

$$H_1|_{\theta=0,\pi} = H_3|_{\theta=0,\pi} = H_6|_{\theta=0,\pi} = \Phi_2|_{\theta=0,\pi} = 0, \quad \partial_\theta H_2|_{\theta=0,\pi} = \partial_\theta H_4|_{\theta=0,\pi} = \partial_\theta \Phi_1|_{\theta=0,\pi} = 0. \quad (47)$$

However, as discussed for the first time in [44] for the GG model, there exists also a different type of solutions of the second order Euler-Lagrange equations, which are not stable and represent saddle points of the energy, rather than absolute minima. In the absence of an electric field, they correspond to magnetic dipoles. A systematic discussion of the properties of these solutions in GG model is given in Ref. [45] (note that these are non-BPS configurations and no exact solution is known in this case). For example, the

magnetic charge measured at infinity vanishes, despite the existence locally of a nonzero density. This, in the presence of an electric charge, results in nonzero angular momentum. The electrically charged, spinning version of the dipole solutions are studied in [12,13].

The magnetic dipole solutions satisfy a different set of boundary conditions at infinity than Eqs. (45), (46), (47). These boundary conditions are found by taking $m = 2$ in the general expressions Eqs. (36), (38)

$$H_1|_{r=\infty} = H_3|_{r=\infty} = 0, \quad H_2|_{r=\infty} = H_4|_{r=\infty} = -1, \\ H_5|_{r=\infty} = V_0 \cos \theta, \quad H_6|_{r=\infty} = V_0 \sin \theta, \quad \Phi_1|_{r=\infty} = \cos \theta, \quad \Phi_2|_{r=\infty} = \sin \theta, \quad (48)$$

(with $V_0 < 1$ again). The other boundary conditions are

$$H_1|_{r=0} = H_3|_{r=0} = 0, \quad H_2|_{r=0} = H_4|_{r=0} = 1, \\ (\cos \theta \partial_r H_5 - \sin \theta \partial_r H_6)|_{r=0} = 0, \quad (\sin \theta H_5 + \cos \theta H_6)|_{r=0} = 0, \\ (\cos \theta \partial_r \Phi_1 - \sin \theta \partial_r \Phi_2)|_{r=0} = 0, \quad (\sin \theta \Phi_1 + \cos \theta \Phi_2)|_{r=0} = 0, \quad (49)$$

at the origin, and

$$H_1|_{\theta=0,\pi} = H_3|_{\theta=0,\pi} = H_6|_{\theta=0,\pi} = \Phi_2|_{\theta=0,\pi} = 0, \quad \partial_\theta H_2|_{\theta=0,\pi} = \partial_\theta H_4|_{\theta=0,\pi} = \partial_\theta \Phi_1|_{\theta=0,\pi} = 0, \quad (50)$$

on the symmetry axis.

1. $n = 1$ results: Spherically symmetric generalized dyons

The spherically symmetric solutions are found by taking

$$H_1 = H_3 = H_6 = \Phi_2 = 0, \quad H_2 = H_4 = w(r), \quad H_5 = u(r), \quad \Phi_1 = h(r) \quad (51)$$

in the general axially symmetric parametrization Eqs. (34), (35) and have a winding number $n = 1$. The boundary conditions in this case can be read from Eq. (45) together with Eq. (46).

The one-dimensional reduced Lagrangian of this system is given by the sum of the $p = 1$ and $p = 2$ terms,

$$\mathcal{L}_{\text{YMH}}^{(1)} = -\frac{1}{4}r^2 \left[2\left(\frac{w'}{r}\right)^2 + \frac{1}{r^4}(1-w^2)^2 \right] - r^2 \left[(\eta^2 h'^2 - u'^2) + 2\left(\frac{w}{r}\right)^2 (\eta^2 h^2 - u^2) \right], \quad (52)$$

and

$$\begin{aligned} L_{\text{YMH}}^{(2)} = & 6\{\lambda_0([(1-w^2)u']^2 - \lambda_1\eta^2([(1-w^2)h']^2)\} \\ & - 12\lambda_2\eta^4\{2([(1-h^2)w']^2 + r^{-2}([(1-w^2)(1-h^2) + 2w^2h^2])^2)\} \\ & + 12\eta^4r^2(1-h^2)^2\{4\lambda_2[u'^2 + 2r^{-2}w^2u^2] - 3\lambda_3\eta^2[h'^2 + 2r^{-2}w^2h^2]\} \\ & - 54\lambda_4\eta^8r^2(1-h^2)^4. \end{aligned} \quad (53)$$

The first of these, Eq. (52), supports the usual Julia-Zee dyon, while the second one, Eq. (53), supports the next excited Julia-Zee dyon. Here, we have analyzed the dyon solutions of a combination of these systems.

The system of three nonlinear coupled differential equations for the functions w , h , and u , subject to the boundary conditions described above, was solved by using the software package COLSYS developed by Ascher, Christiansen, and Russell [46]. This solver employs a collocation method for boundary-value ordinary differential equations and a damped Newton method of quasilinearization.

We have studied in a systematic way the solutions of the $(p = 1) + (p = 2)$ YMH model by varying the parameters λ_i which enter the $p = 2$ reduced Lagrangian Eq. (53). Here, however, we exhibit in Figs. 1 and 2 the results of the numerical integration for two subcases of main interest only. In both cases, we take the GG Lagrangian plus some terms in the $p = 2$ Lagrangian. Namely, we consider what we refer to as a *type (I) model* with $\lambda_0 \neq 0$, $\lambda_a = 0$ ($a = 1, \dots, 4$), and otherwise as a *type (II) model*, with $\lambda_3 \neq 0$, $\lambda_0 = \lambda_1 = \lambda_2 = \lambda_4 = 0$. The first model captures the effects of F^4 corrections originating in the gauge fields, while the type (II) model involves corrections from gauged scalar fields only. Note that the constants λ_0 and λ_3 are made dimensionless by applying a suitable scaling.

In both cases, the properties of the solutions are similar to those of their dyonic axially symmetric generalizations, and thus we shall not discuss them further here.

2. Axially symmetric solutions

We turn now to the case of configurations with a nontrivial dependence of both r and θ . In our approach,

the solutions to the field equations are found by solving numerically a set of eight nonlinear coupled partial differential equations for the functions H_i , Φ_i with the boundary conditions given above. The numerical calculations were performed with the help of the package FIDISOL-CADSOL [47], based on the Newton-Raphson iterative procedure, in which case the known solutions of the GG model provide the initial guess.

However, the $(p = 1) + (p = 2)$ YMH equations with all $\lambda_i \neq 0$ are truly formidable (with some equations containing up to 500 terms), so we eschew the general case. Instead, we restrict our study to two subcases of main interest mentioned above for the spherically symmetric limit of the model.

The axially symmetric generalized dyons reported in this work have a winding number $n = 2$. However, we have constructed several solutions with $n = 3, 4$; thus they are expected to exist for any n . The considered generalized dipoles have a winding number $n = 1$. In practice, we have chosen to fix the value of V_0 and to compute the mass and electric charge of families of solutions by varying λ_0 and λ_3 , respectively. Some numerical output is shown in Figs. 3–6.

Our results can be summarized as follows:

- (i) First, all known dyon and dipole solutions possess generalizations with $p = 2$ terms. However, with $\lambda_i \neq 0$, the generalized dyon solutions satisfy the second-order Euler–Lagrange equations and not the first-order Bogomol’nyi equations as in the pure $p = 1$ case. As a result, the solutions always possess a nonvanishing angular momentum density, $T_\phi^t \neq 0$ (despite the fact that the total angular momentum of the generalized dyons vanishes).

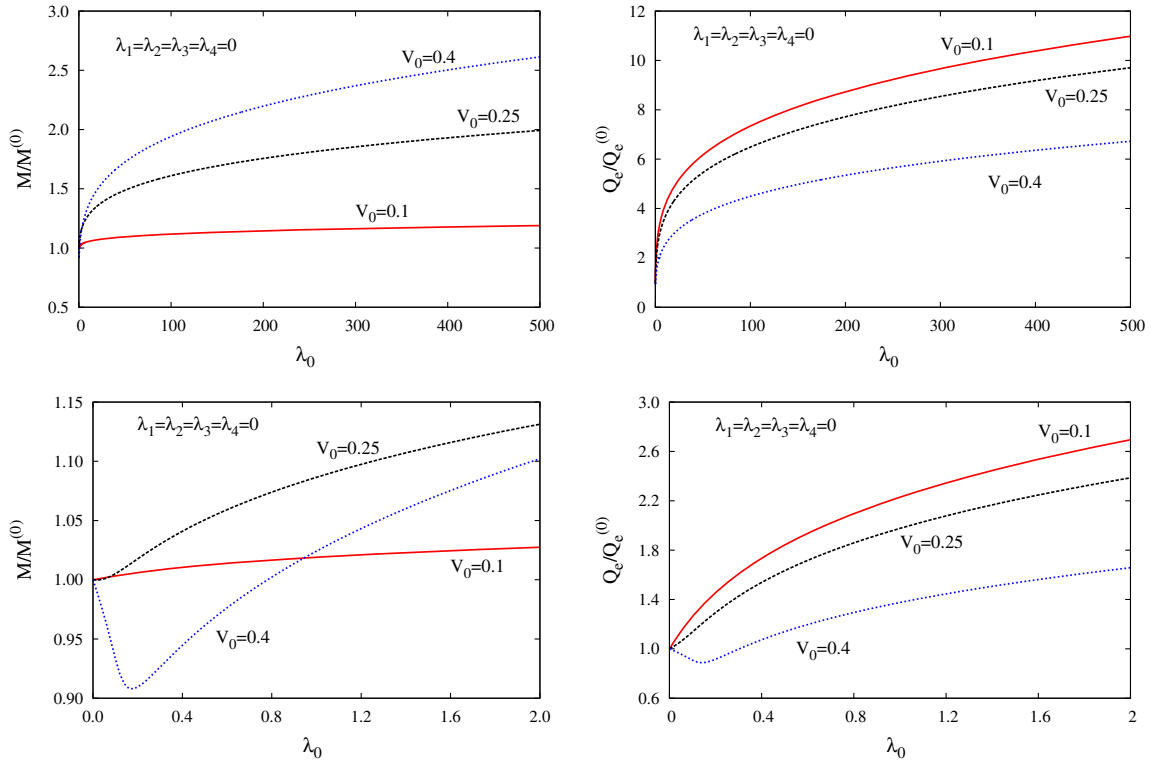


FIG. 1 (color online). The mass M and the electric charge Q_e are shown as a function of the coupling constant λ_0 for type (I) spherically symmetric generalized dyon solutions. Several values of the electric potential V_0 are considered. The bottom panel is a zoom-in plot of the top one. Here and in Figs. 2–6, M and Q_e are normalized with respect to the corresponding solutions in the Georgi-Glashow model with the same V_0 .

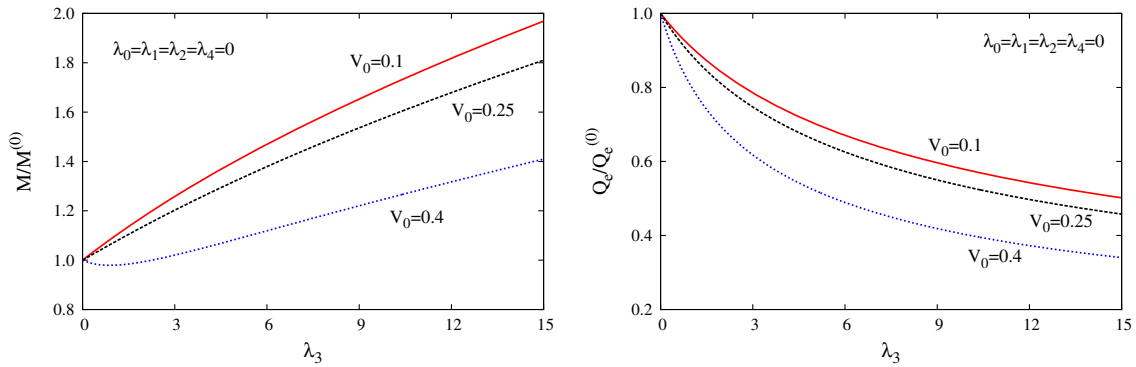


FIG. 2 (color online). The mass M and the electric charge Q_e are shown as a function of the coupling constant λ_3 for type (II) spherically symmetric generalized dyon solutions for several values of the electric potential V_0 .

- (ii) For all solutions, the profiles of the functions H_i and Φ_i look similar to the corresponding ones in the GG limit. The same holds for the distribution of the energy and angular momentum densities. For generalized dyons, the energy density has a strong peak along the ρ axis, and it decreases monotonically along the symmetry axis. Equal density contours reveal a toruslike shape of the configurations, the tori being localized in the equatorial plane. For the $n = 1$ generalized dipoles, the energy density

always possesses maxima on the positive and negative z axis at the locations of the monopole and antimonopole and a saddle point at the origin. Equal density contours consist in two tori on the symmetry axis.

- (iii) Another property of GG model that persists in the presence of $p = 2$ corrections is that found by Houston and O’Raifeartaigh [48]: any regular axially symmetric magnetic charge distribution can be located only at isolated points situated on the axis of

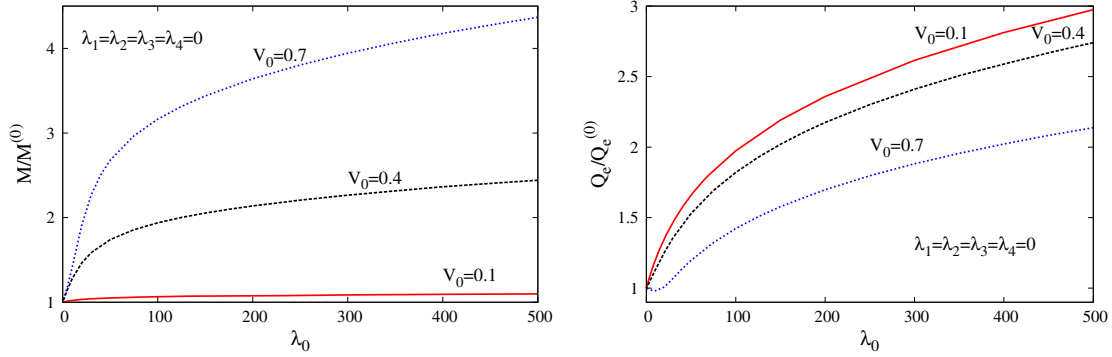


FIG. 3 (color online). The mass M and the electric charge Q_e are shown as a function of the coupling constant λ_0 for type (I) axially symmetric generalized dyon solutions. Several values of the electric potential V_0 are considered. The solutions here and in Figure 4 have a winding number $n = 2$.

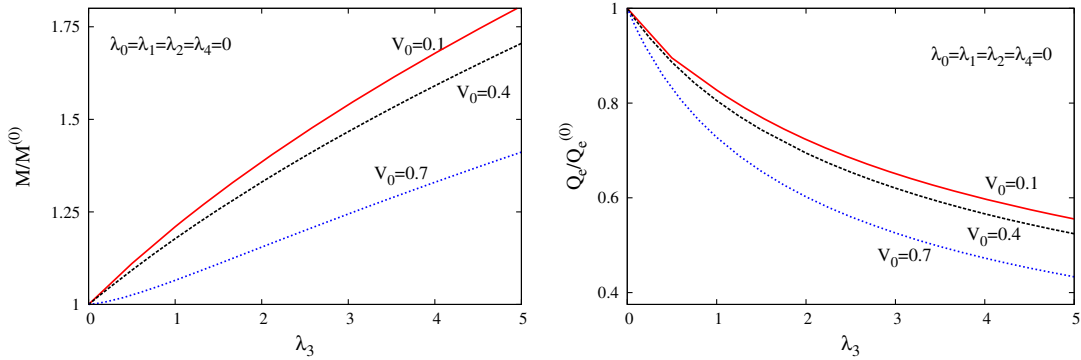


FIG. 4 (color online). The mass M and the electric charge Q_e are shown as a function of the coupling constant λ_3 for type (II) axially symmetric generalized dyon solutions and three values of the electric potential V_0 .

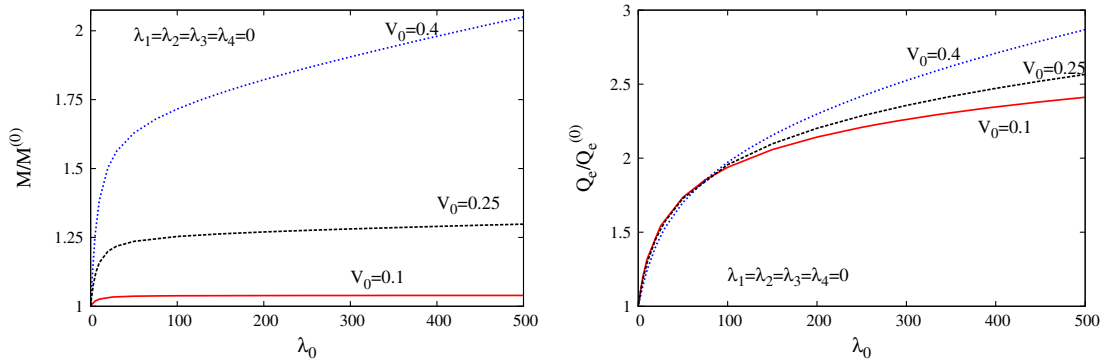


FIG. 5 (color online). Same as Fig. 3 for electrically charged, generalized dipole solutions. Here and in Fig. 6 the winding number is $n = 1$.

symmetry, with equal and opposite values of the charge at alternate points. In particular, if only one sign of the magnetic charge is allowed (i.e. for generalized dyons), all the magnetic charge is concentrated at the origin, where the Higgs field vanishes.

- (iv) No upper bounds seem to exist on the values of the coupling constants λ_0 , λ_3 . As seen in Figs. 1 and 3,

adding an F^4 term to the GG model [a type (I) model] and taking large enough values of λ_0 , increases the mass and the electric charge of the solutions (this feature holds for any value of the electric potential V_0). However, rather unexpectedly, the behavior is different for small λ_0 and large V_0 , with the existence of a minimum for both M and Q_e below the values found in the GG model (see Fig. 1).

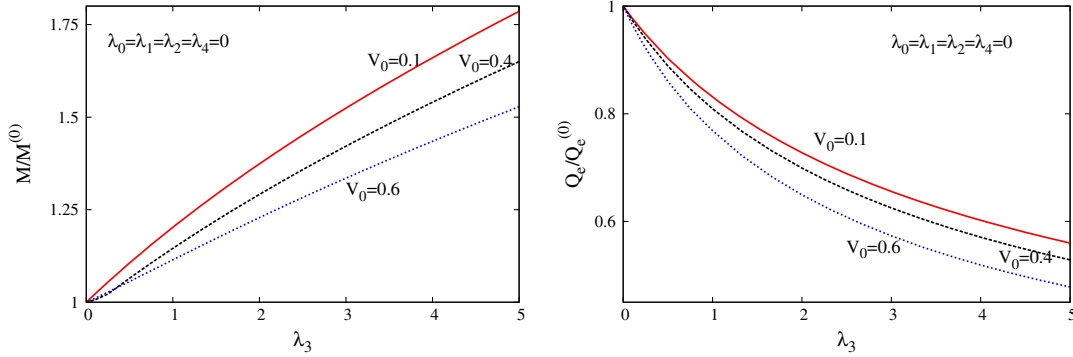


FIG. 6 (color online). Same as Fig. 4 for electrically charged, generalized dipole solutions.

At the same time, while the mass increases with λ_3 for the type (II) model, the electric charge decreases (this holds for both generalized dyons and generalized dipoles; see Figs. 4 and 6).

More complex configurations representing chains of m monopoles and antimonopoles are known to exist for the case of a GG model [40]. We expect these solutions to possess also generalizations within the $(p = 1) + (p = 2)$ YMH model.

VI. FURTHER REMARKS

The main purpose of this work was to address the question on how general the relation between angular momentum and electric and magnetic charges is, originally derived in [11] for the GG model. We have considered various generalizations of the GG model and have found that the conjecture in [11] that *a nonvanishing total angular momentum is incompatible with a net magnetic charge* remains valid. Perhaps most prominently, we have considered the correction of the GG model by adding a *fourth-order* Yang-Mills-Higgs density, which also does not result in an angular momentum for the corresponding axially symmetric generalized dyon. There is no question that this qualitative conclusion will remain valid when $2p$ th order YMH terms are introduced. In addition to these, we considered the Born-Infeld-Higgs system, as well as a model [35] where a peculiar parity-charge violating term is added to the GG Lagrangian. In both cases precisely the same conclusion was arrived at. This result is independent of the dynamics, at least for the models considered here, which we believe is exhaustive. It relies on the asymptotic behaviors of the gauge potential and Higgs field at infinity. The mechanism behind this result turns out to be that the presence in the Lagrangian of the usual “quadratic kinetic” term $D_\mu\Phi D^\mu\Phi$ of the Higgs field enforces the same asymptotic behavior at infinity as in that of the GG model.

Whether there is a possibility of circumventing this obstacle seems unlikely. Indeed, it is possible to exclude the usual Higgs “quadratic kinetic” term from the Lagrangian by choosing e.g. to work exclusively with an (unphysical) model consisting of the usual F^2 YM term plus the $p = 2$

Lagrangian Eq. (22), i.e. in the absence of the usual $D_\mu\Phi D^\mu\Phi$ Higgs kinetic term. We have constructed such solutions but only in the spherically symmetric limit. In that case one finds solutions with essentially different (non-standard) asymptotics of the magnetic gauge potential. Notably in this case the asymptotic YM connection $w(r)$ decays in a manner that it does not describe a magnetic monopole but rather some other unipole like e.g. the Skyrme hedgehog. From this, one might infer that the angular momentum surface integral (for the corresponding axially symmetric solutions when they are found) might conceivably not vanish. This possibility is very unlikely because a nonvanishing angular momentum is known to be concurrent with a nonvanishing electric field, and, when this (nonstandard) spherically symmetric soliton is charged with an electric field in the manner of Julia and Zee, the asymptotics revert to the standard asymptotics of the “magnetic monopole” type, that precludes nonvanishing global angular momentum. We will elaborate on the details elsewhere and have given here a brief description in the Appendix to support this claim.

Concerning various possible generalizations of the results in this work, let us mention first that the inclusion of the gravity effects does not change the general relations in [11] between the angular momentum and magnetic and electric charges, as long as the configurations do not possess an event horizon.¹⁰ The situation changes for black hole solutions; for example, the total angular momentum of a dyonic black hole solution is nonzero, due to the contribution of the angular momentum [49]. However, a discussion of these aspects is beyond the scope of the present work.

An interesting version of the problem considered in this work is the pure YM limit, i.e. no Higgs field. In the pure F^2 -YM theory, a number of well-known results forbids the existence of particlelike soliton solutions. In order for such configurations to exist, one has to couple the model to

¹⁰Moreover, the same results hold for anti-de Sitter asymptotics of the spacetime. The case of a positive cosmological constant has not been yet considered in the literature.

gravity [4] (or, simpler, to a dilaton field [50]). Interestingly, as argued in [11], in contrast to other solitons, these solutions do not possess spinning generalizations. However, considering a more general Lagrangian of the YM field (for example with an extra F^4 term or working directly in a Born-Infeld theory) may give hope to circumventing this result (note that a non-Abelian Born-Infeld theory allows for YM solitons without coupling to any other field [15]). Unfortunately, this is not the case, the mechanism put forward in [11] for a F^2 theory being still valid in this case. Again, the problem resides in the asymptotic behavior of the gauge potentials: since the F^2 contribution would dominate in the field equations as $r \rightarrow \infty$, the electric components of the gauge field act like an isotriplet Higgs field with negative metric, and by themselves cause some of the magnetic components to oscillate rather than decrease exponentially, which would lead to delocalized configurations. Therefore, we are forced to take a vanishing asymptotic value of the electric gauge potential, $V_0 = 0$. However, from Eq. (20) this implies $A_i \equiv 0$ (since $\mathcal{E} \geq 0$), i.e. a static configuration.

Finally, let us address the question of spinning solutions in the standard model. The results in [51,52] show that the well-known Klinkhamer-Manton sphalerons possess spinning generalizations. Moreover, for a nonzero mixing angle θ_w , the angular momentum of a spinning sphaleron equals the electric charge [51,52] (note the analogy with the monopole-antimonopole pair). However, the Lagrangian of the Weinberg-Salam model can be supplemented with higher derivative corrections, the simplest one consisting in Skyrme-like terms of the Higgs field plus F^4 terms of the gauge fields. Based on the results in this work, we conjecture that the qualitative results in [51,52] remain always valid and the angular momentum of a spinning sphaleron always equals the electric charge.

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APPENDIX: SOLITONS OF THE F^2 PLUS $p = 2$ YMH SYSTEM IN $3 + 1$ DIMENSIONS

In this Appendix we consider a model consisting of the usual F^2 Yang-Mills term plus the pure $p = 2$ YMH system in $3 + 1$ dimensions. Our original aim was to find solutions with a more general asymptotic behavior of the gauge potentials, in the hope of circumventing the ban on spinning YMH monopole solitons. As this attempt gave a final negative result, the analysis was relegated to the Appendix, in support of our claim that topologically stable YMH monopoles cannot spin.

It is important in this context to distinguish between a topologically stable monopole (or unipole) and a topologically stable ‘‘magnetic monopole.’’ The hedgehog of the Skyrme model is a monopole centered at the origin. The ’t Hooft-Polyakov hedgehog on the other hand is a monopole, with the additional property that its asymptotic gauge field is a Dirac-Yang $U(1)$ Maxwell field. In this case the magnetic gauge function $w(r)$ vanishes asymptotically, i.e. $w(r) = 0$ as $r \rightarrow \infty$. By contrast, the solutions to the system

$$\mathcal{L} = -\frac{1}{4}\alpha\text{Tr}(F_{\mu\nu}F^{\mu\nu}) + \mathcal{L}^{(2)}, \quad (\text{A1})$$

with $\mathcal{L}^{(2)}$ given by Eq. (22), have $\lim_{r \rightarrow \infty} w(r) \neq 0$.

Before discussing this case, let us consider first the pure $p = 2$ YMH magnetic system (i.e. without a F^2 term, $\alpha = 0$). The corresponding energy density functional is found by taking in Eq. (53) $u(r) = 0$ together with $\lambda_0 = 0$, $\lambda_1 = \lambda_2 = \lambda_3 = \lambda_4 = 1$, and reads (up to some unimportant overall numerical factor)

$$\begin{aligned} H_{\text{mag}}^{(2)} &= \eta^2([(1-w^2)h]')^2 + 3^2\eta^8 r^2(1-h^2)^4 \\ &\quad + 2^2\eta^4([(1-h^2)w]')^2 + 3^2 \cdot 2^4\eta^6(1-h^2)^2 w^2 h^2 \\ &\quad + 3^2 \cdot 2^3\eta^6 r^2(1-h^2)^2 h'^2 \\ &\quad + 2\eta^4 r^{-2}[(1-h^2)(1-w^2) + 2w^2 h^2]^2, \end{aligned} \quad (\text{A2})$$

being bounded from below by the topological charge density [25]

$$\rho = 2\eta^5 \frac{d}{dr} \left\{ \left(h - \frac{2}{3}h^3 + \frac{1}{5}h^5 \right) - (1-h^2)^2 w^2 h \right\}. \quad (\text{A3})$$

However, this bound is never achieved, since the Bogomol’nyi equations of the pure $p = 2$ model

$$\begin{aligned} \eta r^{-1}[(1-w^2)h]' &= \pm 3\eta^4 r(1-h^2)^2, \\ \eta^2[(1-h^2)w]' &= \mp 3\eta^3(1-h^2)wh, \\ 3\eta^3 r(1-h^2)h' &= \pm \eta^2 r^{-1}[(1-h^2)(1-w^2) + 2w^2 h^2] \end{aligned} \quad (\text{A4})$$

are overdetermined [43].

Returning to the case of the model

$$H = \frac{1}{4}\alpha r^2 \left[2 \left(\frac{w'}{r} \right)^2 + \frac{1}{r^4} (1-w^2)^2 \right] + H_{\text{mag}}^{(2)},$$

with arbitrary λ_i [and $u(r) = 0$], one notices that the regularity of the solutions imposes the same boundary conditions at $r = 0$ as in the $p = 1$ case, i.e. $w(0) = 1$ and $h(0) = 0$. However, the far field asymptotics can be different, the magnetic gauge potential $w(r)$ possessing

another asymptotic value compatible with finite energy requirement, apart from the standard one $w(\infty) = 0$. We have found numerical evidence for the existence of solutions smoothly interpolating between $w(0) = 1$ and $h(0) = 0$ and

$$h(r) \rightarrow 1, \quad w(r) \rightarrow \frac{1}{\sqrt{1+192\eta^4\lambda_2/\alpha}} \text{ as } r \rightarrow \infty. \quad (\text{A5})$$

Their total mass is finite, possessing a nontrivial dependence on the parameters λ_i .

The situation is, however, different when looking for electric generalizations of these configurations. It turns out that no solutions compatible with the finite energy

requirements can be found for the asymptotics Eq. (A5). This can be understood as follows. First, the energy density possesses a u^2w^2 term, which originates in the F^2 part of the Lagrangian. This term should decay faster than $1/r$ as $r \rightarrow \infty$; thus a value $w(\infty) \neq 0$ results in $u(\infty) = V_0 = 0$. However, this implies directly a vanishing electric potential, as seen from the relation Eq. (41), which still holds in this case. Thus we conclude that a nonzero electric potential is not compatible with nonstandard asymptotics of the magnetic gauge potential in the F^2 plus $p = 2$ YMH system.

On general grounds, we expect a similar result to hold as well in the axially symmetric case.

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